

18 Downstream development during South African cut-off low  
19 development

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28 **Abstract**

Using 39 years of ECMWF reanalysis data, an established energetics framework and simple composite analysis this study has shown that South African cut-off low (COL) pressure systems are preceded by downstream development of a baroclinic wave. The upstream eddy kinetic energy, which is associated with the midlatitude jet streak, develops and reaches its maximum before the formation of the closed COL cyclonic circulation. The downstream eddy kinetic energy centre maximises at the point which the closed circulation forms. The upstream eddy kinetic energy centre grows from baroclinic conversion from eddy available potential energy to eddy kinetic energy, whilst the latter grows by receiving energy by means of ageostrophic geopotential fluxes that transport eddy kinetic energy in a north-eastward direction from the upstream centre. These ageostrophic geopotential fluxes are induced, increased in magnitude and directed by processes associated with RWB on the midlatitude dynamical tropopause. and so the downstream energy transfer connects South African COLs to midlatitude processes. The study has further shown that the baroclinic kinetic energy configuration previously associated with wet seasons over South Africa is consistent with times when COLs forms over the country. This study shows further that these two branches are linked by the ageostrophic geopotential fluxes, for COLs that occur in the western half of South Africa.

29 *Keywords:* Downstream development, cut-off lows, eddy kinetic energy

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## 5 Highlights

### 6 **Downstream development during South African cut-off low development**

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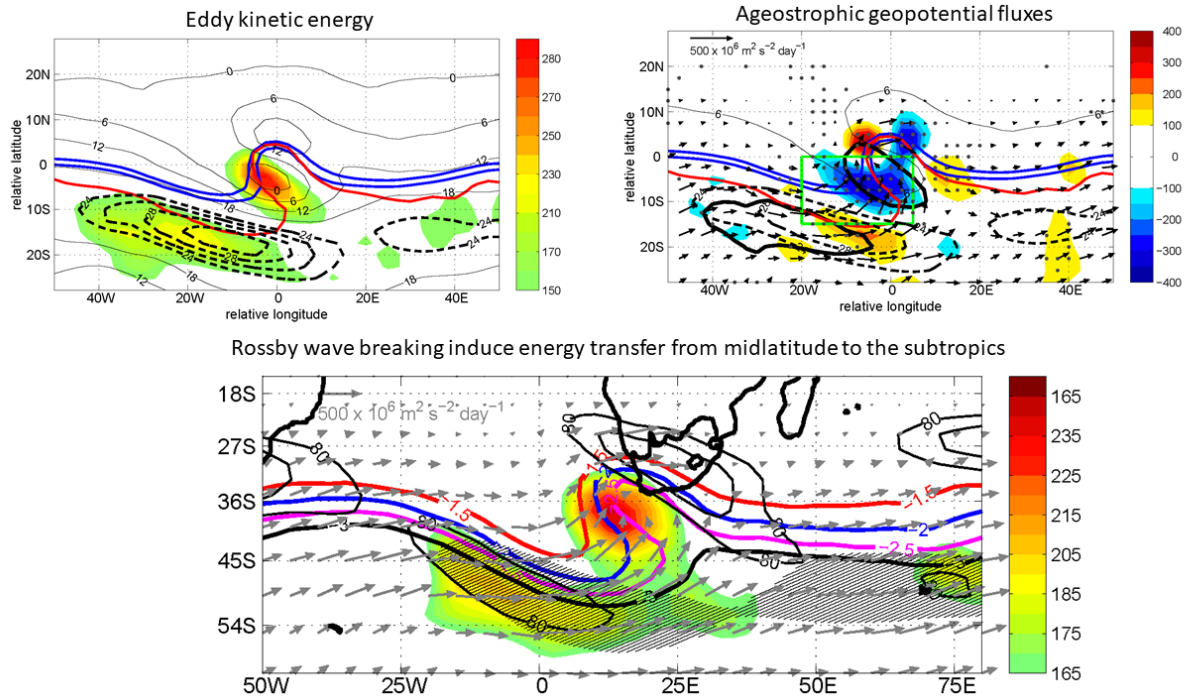
- 9     • The study has shown that South African cut-off low (COL) pressure systems  
10       are preceded by downstream development.
  
- 11     • Downstream development in the context of the COLs provides a framework to  
12       show that the eddy kinetic energy associated with COL formation is transferred  
13       from the midlatitudes, is not converted from eddy available potential energy via  
14       baroclinic conversion.
  
- 15     • Breaking Rossby waves in the midlatitudes induce and enhance the ageostrophic  
16       geopotential fluxes that make the transfer of energy from the midlatitudes to  
17       the subtropics possible.

1 Graphical Abstract

2 Downstream development during South African cut-off low development

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## 30 1. Introduction

31 A cut-off low (COL) pressure system is a closed and cold cored upper tropospheric  
32 cyclonic circulation that has been detached from the westerlies (Palmén and Newman,  
33 1969; Pinheiro et al., 2017, 2019) on the anticyclonic barotropic shear side of the jet  
34 core (Thorncroft et al., 1993). This closed circulation is induced by high potential  
35 vorticity (PV) anomalies (Hoskins et al., 1985) that may be brought about by the  
36 equatorward isentropic transport of stratospheric air by means of Rossby wave break-  
37 ing (RWB) processes (Ndarana and Waugh, 2010; Reyers and Shao, 2019; Barnes et  
38 al., 2000a). McIntyre and Palmer (1983) defined RWB as the rapid and irreversible  
39 deformation of PV contours that turn back on themselves such that the meridional  
40 gradient of the PV is reversed. Such a reversal is most clearly seen when PV is pre-  
41 sented on isentropic surfaces (Hoskins et al., 1985). Other studies have defined RWB  
42 as overtuning isentropic contours on iso-PV surfaces (e.g. Berrisford et al., 2007) and  
43 as the deformation of absolute vorticity on isobaric surfaces (Barnes and Hartmann,  
44 2012). The fact that the closed COL circulation is induced by these anomalies means  
45 that its dynamical onset is an upper tropospheric process, but may extend to the  
46 lower levels of the atmosphere (Barnes et al., 2020a, 2020b; Portmann et al., 2020)  
47 and may induce surface cyclogenesis (Portman et al., 2020). When COLs extend to  
48 the surface they may lead to storm surge that can potentially cause extensive damage  
49 along the coastal areas in the South African domain (Barnes et al., 2020b).

50 COLs in the South African domain have been extensively studied using reanalysis  
51 products (e.g. Singleton and Reason, 2007a; Favre et al., 2012, 2013) and using nu-  
52 merical models e.g. Singleton and Reason, 2006, 2007a), and have clearly been shown  
53 to bring rainfall to different parts of the country (Molekwa et al., 2014; Engelbrecht  
54 et al., 2015; Omar and Obiodun, 2020). The typically heavy rainfall associated with  
55 COLs implies that east of the upper air trough axis substantial upward motion exists.  
56 This vertical uplift is a well known dynamical characteristic of a fully developed baro-  
57 clinic weather system, with convergence at the surface and divergence aloft. Such a  
58 system dissipates when the upper level disturbance catches up with the one at the  
59 surface so that its vertical structure becomes equivalent barotropic. Therefore, when

60 COLs reach South Africa as rainbearing systems; they are usually already matured  
61 synoptic weather systems, at least from a dynamic meteorology point of view. To  
62 support this notion, given the fact that the COLs are preceded by RWB (Ndarana  
63 and Waugh, 2010; Reyers and Shao, 2019), as already mentioned, Favre et al. (2013)  
64 and Omar and Obiodun (2020) consistently note that they develop from unstable  
65 Rossby waves and, due to this instability, these waves eventually evolve into non-  
66 linear regimes and break. This breaking is actually a dissipative process (Methven  
67 et al., 2005) during baroclinic life cycles. When viewed in the context of an ide-  
68 alised experiments (e.g. Thorncroft et al., 1993; Kunz et al., 2009; Kunkel et al.,  
69 2011), irreversible deformation of the PV contours that signals RWB occurs during  
70 the downturn of eddy kinetic energy ( $K_e$ ), when the vertical propagation of wave  
71 activity has ceased and its meridional component dominates. This happens during  
72 the barotropic conversion stage and the  $K_e$  decreases, as it is converted into mean  
73 kinetic energy.

74 Approaches that have been taken to study COLs in the South African domain  
75 have so far focused on processes from the formation of the closed circulation and  
76 subsequent impacts, even though linkages with low frequency atmospheric oscillations  
77 and remote drivers have been considered (Singleton and Reason, 2007a; Favre et al.,  
78 2012) and boundary conditions, Singleton and Reason (2006, 2007b) simulated COLs  
79 over South Africa using MM5 and analysed the contribution of SSTs, surface latent  
80 fluxes and topography to the evolution of the COL and its associated rainfall. A key  
81 outstanding question that warrants analysis is what happens dynamically on days  
82 leading up to the formation of the closed circulation.

83 The connection between COL pressure systems and RWB then suggests that the  
84 former may also be viewed in the context of a baroclinic wave life cycle using the  
85 energetics framework. For example, Gan and Piva (2013, 2016) employed the Or-  
86 lanski and Katzfey (1991) system of local energy equations to study the evolution of  
87 southeastern Pacific COL pressure systems. They showed that COLs were associated  
88 with two  $K_e$  centres of differing spatial extent, with the larger one coinciding with the  
89 jet streak identified in Ndarana and Waugh (2010), Reyers and Shao (2019). Aspects

90 of the evolution of the jet streak were explained in Ndarana et al. (2020). Based on  
91 the Gan and Piva (2013) case study,  $K_e$  is then transported from the jet stream via  
92 ageostrophic geopotential fluxes into the closed circulation region. However whether  
93 this is the case in the South African domain and the cause of the increase in the  
94 strength (and perhaps even the onset) of this energy transport in South African COL  
95 pressure systems remains an open question.

96 Tennant and Reason (2005) showed that kinetic energy plays a role in regulating  
97 rainfall variability over South Africa. They found that during the wet seasons, the  
98 baroclinic kinetic energy breaks into two branches, with one branch located in the  
99 subtropics and another in the midlatitudes. They considered these energetics issues of  
100 South African rainfall at the seasonal time scale, thus providing a framework within  
101 which the questions raised above could be considered. One of the hypotheses that  
102 the current study explores is that there could be a midlatitude/subtropical baroclinic  
103 kinetic energy transfer that is characterised by the ageostrophic geopotential fluxes,  
104 even though there is no theoretical framework that we are aware of that links the  
105 Orlanski and Katzefey (1991) and Wiin-Nielson (1962) energetics frameworks (the  
106 latter of which was used in Tennant and Reason (2005)).

107 Given the above, the overall aim of this study is to establish the dynamical pro-  
108 cesses that take place during the life cycle of South African COLs, assuming that  
109 the onset of this life cycle occurs before the closed circulation forms. This study will  
110 also link the energetics of COLs to the Tennant and Reason (2005) result. As these  
111 are high impact weather systems, such an analysis has the potential for contributing  
112 to improving their predictability, particularly at the medium range forecasting time  
113 scale. The rest of the paper is organised as follows: In the next section the data and  
114 methods are presented. The results Section 3 is divided into subsections that dis-  
115 cuss aspects of the dynamical evolution of COLs. Section 4 presents the concluding  
116 remarks.

## 117 2. Data and methods

### 118 2.1. Data

119 As in Barnes et al. (2020a) and Ndarana et al. (2020), we identify COLs on the 500  
120 hPa isobaric level using the European Centre for Medium-Range Weather Forecasting  
121 (ECMWF) Reanalysis Interim (ERA-I; Dee et al., 2011). In this diagnostics study, the  
122 geopotential height, the three components of the wind and temperature fields were  
123 extracted for calculating the mathematical quantities described below. The study  
124 covers the years 1979 to 2018 because prior to 1979, reanalysis datasets were unreliable  
125 in the Southern Hemisphere (Tennant, 2004) as the data would pre-date the advent  
126 of satellite observations and the assimilation thereof into global data assimilation  
127 systems.

### 128 2.2. Methods

#### 129 2.2.1. COL Detection and composite analysis

130 The method for identifying COLs and calculating the composites have been de-  
131 scribed in Barnes et al. (2020a) and Ndarana et al. (2020), and only a brief descrip-  
132 tion of it is outlined here. It is an 8-step algorithm that starts by identifying closed  
133 contours on the 500 hPa geopotential height field. Centres of these closed contours  
134 that are north of  $15^{\circ}\text{S}$  and south of  $50^{\circ}\text{S}$  are excluded. To ensure that the circulation  
135 around the contour is cyclonic, geopotential heights at the centres are then assumed  
136 to be lower than 6 of the immediately surrounding grid points. Once this condition  
137 has been been complied with, it is further required that the zonal component of the  
138 wind south of the centre be negative to ensure that a cut-off has occurred.

139 As a fifth step, the cold core condition is imposed, after which all concentric  
140 contours are grouped together. These are defined as all closed contours whose centres  
141 are within a  $10^{\circ} \times 10^{\circ}$  grid box. The centre with the lowest 500 hPa geopotential height  
142 value is considered to be the final COL centre point. In the final step, COL centres  
143 that lie within 1000 km from one another on consecutive time steps are considered to  
144 belong to the same evolving COL. As in Singleton and Reason (2007a), we retain only  
145 those systems that have a 24 hour duration or longer within the domain bounded by

146 10 - 40°E longitude and 20 - 40°S latitude.

147 As in Ndarana et al. (2020), we use the centre of the COL as a reference point  
148 when calculating the composite means. Firstly all the fields bounded by a 100°  
149 longitude - 50° latitude box, with the middle most point of the box being the COL  
150 point are extracted from each field. We then bring these together, so that the middle  
151 most points of the fields coincide and then calculate simple means of the fields. No  
152 assumption is made except the fact that a closed circulation has formed. It thus  
153 follows that the fields that emerge are representative of the COLs. The statistical  
154 significance is calculated using the approach of Brown and Hall (1999).

### 155 *2.3. Diagnostics*

To generate the local energetics fields that evolve in space and time, we assume a basic state flow defined by the 31-day time mean (Orlanski and Katzfey, 1991; Lackmann et al., 1999; Danielson et al., 2006) centred on the day of the COL events and the perturbations are the deviations from that time mean. The choice of the 31-day time mean basic state flow leads to simpler equations and Lackmann et al. (1999) suggested that low frequency basic state flow does not significantly influence the energy budget analysis results. The total variables are then decomposed as follows:

$$\mathbf{U} = \mathbf{U}_m + \mathbf{u} \quad (1)$$

$$\mathbf{V} = \mathbf{V}_m + \mathbf{v} \quad (2)$$

$$\Phi = \Phi_m + \phi \quad (3)$$

$$\Theta = \Theta_m + \theta \quad (4)$$

156 where the capital letters/Greek symbols with no subscript  $m$  and with the subscript  
157  $m$  represent the total and time mean variables, respectively. The lowercase symbols  
158 represent the perturbation fields and  $\mathbf{U} = U\mathbf{i} + V\mathbf{j} + \omega\mathbf{k}$  is the three dimensional  
159 velocity flow and  $\mathbf{V} = U\mathbf{i} + V\mathbf{j}$  is the horizontal flow on isobaric surfaces. The  
160 symbols  $\Phi$  and  $\Theta$  are the geopotential and potential temperature, respectively.

We employ the flux form of the Orlanski and Katzfey (1991) and Danielson et al.



(2006) eddy available potential energy ( $P_e$ ) equation given by

$$\partial_t P_e = -\nabla_p \cdot (\mathbf{V}P_e) - \partial_p(\omega P_e) + \omega\alpha + \frac{\alpha_m}{2\Theta_m} \frac{1}{d\widehat{\Theta}/dp} (\mathbf{v}\theta \cdot \nabla_p \Theta_m) + S \quad (5)$$

where  $\alpha$  is the specific volume and the other terms have been defined above. The subscript  $p$  in the gradient operator means that it is evaluated whilst keeping pressure constant. As in Orlanski and Sheldon (1993), the correlation between the perturbation temperature and the time mean temperature advection by the eddies as well as other terms that materialised when transforming the advection form of the equation to the flux version of it are all included as non-conservative sources and sinks of  $P_e$ , which are represented by  $S$  in Eq (5). The form of  $P_e$  used in this study is

$$P_e = - \left( \frac{\alpha_m}{2\Theta_m} \frac{1}{d\widehat{\Theta}/dp} \right) \theta^2 \quad (6)$$

161 and the hat over  $\Theta$  represents the horizontal average. The first two terms on the right  
 162 hand side of Eq. 5 are the horizontal and vertical eddy available potential energy flux  
 163 terms, the third term is the baroclinic conversion of  $P_e$  to  $K_e$  and the fourth term  
 164 represents the conversion from mean available potential energy ( $P_m$ ) to  $P_e$

McLay and Martin (2002) derived the flux form of the Orlanski and Katzfey (1991) eddy kinetic equation. Here the zonal and meridional momentum correlation, the vertical flux divergence of the rate of work by aerodynamic stress and the curvature terms are all incorporated in the residual term as was the case above, to obtain

$$\partial_t K_e = -\nabla_p \cdot (\mathbf{V}K_e) - \partial_p(\omega K_e) - \mathbf{v} \cdot \nabla_p \phi + [\mathbf{v} \cdot (\mathbf{u} \cdot \nabla) \mathbf{V}_m] + \text{Residual} \quad (7)$$

165 which is the  $K_e$  equation and  $K_e = \mathbf{v} \cdot \mathbf{v} \times 0.5$

Orlanski and Katzfey (1991) decomposed the  $K_e$  generation term  $-\mathbf{v} \cdot \nabla_p \phi$  as follows

$$-\mathbf{v} \cdot \nabla_p \phi = -\omega\alpha - \nabla_p \cdot (\mathbf{v}\phi)_a - \partial_p(\omega\phi) \quad (8)$$

The first term in Eq. 8 ( $-\omega\alpha$ ) is the same as the third term in Eq. 5 but with opposite sign, thus confirming it as a conversion term between the two energy forms. This term also materises in the contest of the highly idealised two layer model (see Holton and Hakim, 2014) and takes the form  $\overline{\omega_2\psi_T}$ , where  $\psi_T$  is the baroclinic perturbation streamfunction and the overline represents the zonal average over a wavelength of an idealised baroclinic wave. As in Orlanski and Sheldon (1993), we assume a variable  $f$  so that

$$(\mathbf{v}\phi)_a = \mathbf{v}\phi - \mathbf{k} \times \nabla \left( \frac{\phi^2}{2f(y)} \right) \quad (9)$$

166 which is the ageostrophic geopotential flux. It follows then that  $K_e$  is generated  
 167 by two processes, namely (a) baroclinic conversion (first term in Eq. (8)), which is  
 168 caused by vertical eddy heat fluxes and (b) ageostrophic flux convergence (second  
 169 term on Eg. (8)). This system of energy equations was then used by Orlanski and  
 170 Sheldon (1995) to describe downstream development, involving two energy centres,  
 171 one upstream and the other downstream, during which  $K_e$  is moved by means of  
 172 energy fluxes ( $\nabla_p \cdot (\mathbf{V}K_e)$ ), whilst the upstream centre radiates energy downstream  
 173 into the centre to the east of it by means of the ageostrophic geopotential fluxes in  
 174 Eq. (9).

175 In their study of the associations between the global energy cycle and South  
 176 African rainfall, Tennant and Reason (2005) employed a different energetics frame-  
 177 work from the one discussed above. Following Wiin-Nielson (1962), Tennant and  
 178 Reason (2005) decomposed kinetic energy into barotropic and baroclinic compo-  
 179 nents, defined as  $K_{BT} = \mathbf{V}_{BT} \cdot \mathbf{V}_{BT} \times 0.5$  and  $K_{BC} = \mathbf{v}_{BC} \cdot \mathbf{v}_{BC} \times 0.5$ , respectively.  
 180  $\mathbf{V}_{BT}$  is the vertically integrated horizontal flow and  $\mathbf{v}_{BC} = \mathbf{V} - \mathbf{V}_{BT}$ . All the results  
 181 shown below are pressure-weighted vertical average of the diagnostics.

### 182 **3. Results**

#### 183 *3.1. Eddy available potential energy budget*

184 Classical energetics theory (Lorenz, 1955) states that the  $P_m$  in the atmosphere  
 185 is generated by diabatic processes, but converted to  $P_e$ , and then to  $K_e$  by means of

186 baroclinic processes, which are dynamical in nature. This sequence of events dictates  
 187 that we discuss the  $P_e$  budget, using Eq. 5, associated with COLs first and then  
 188 proceed to that of  $K_e$  in the subsequent subsections. Fig. 1 (a) to (e) shows the  
 189 composite evolution of  $P_e$  (see Eq. 6) that is associated with 48-hour COL events  
 190 in the domain, relative to the jet streak (Keyser et al., 1985). This jet streak that  
 191 forms and propagates during the COLs has been reported in several previous studies  
 192 (Ndarana and Waugh, 2010; Reyers and Shao, 2019; Ndarana et al., 2020). Its  
 193 eastward translation is caused by the advection of zonal momentum by the zonal flow,  
 194 whilst the meridional advection of zonal momentum changes its northwest/southeast  
 195 orientation to become more zonal as it passes the closed circulation (Ndarana et al.,  
 196 2020). As the jet streak propagates eastward, the  $P_e$  centre moves along with it, and it  
 197 is always located in the diffluence region of the jet streak. It reaches its maximum at  
 198 the point when the COLs form and subsequently dissipates beyond that point as the  
 199 jet streak propagates further eastward. The co-location of the jet streak and  $P_e$  is the  
 200 first indication that, not only does the former bring about the anticyclonic barotropic  
 201 shear that is necessary for lower stratospheric/upper tropospheric Rossby waves to  
 202 break (Peters and Waugh, 2003; Ndarana and Waugh, 2010; Bowley et al., 2019) as  
 203 indicated by the thick red contour in Fig. 1 (a) to (e) that represents the  $PV = -2$   
 204 PVU ( $1PVU = 10^{-6} \text{ m}^2 \text{ s}^{-1} \text{ K kg}^{-1}$ ), it is also a source region for  $P_e$ . Because this  
 205 PV contour represents the dynamical tropopause (Hoskins, 1991) in the midlatitudes,  
 206 the potential vorticity anomalies that induce the closed COL circulation (Hoskins et  
 207 al., 1985), demonstrate that there is midlatitude stratospheric air presence where the  
 208 South African COLs form.

Note that  $P_e$  has an approximately oval structure and there are areas of  $\partial_t P_e < 0$   
 ( $\partial_t P_e > 0$ ) at its rear (front) ends. This is shown in Fig. 1(f) to (j) by the blue (yellow  
 to brown) shading located in the western (eastern) half of the thick black contour that  
 represents  $P_e = 100 \text{ m}^2 \text{ s}^{-2}$ . It is interesting to note that the  $\partial_t P_e$  field is orientated in  
 the same way as  $\partial_t u$  (cf. Fig. 3 in Ndarana et al., 2020) and temporally and spatially  
 coincides with it. Therefore the eastward movement of  $P_e$  is dynamically coupled to  
 that of the jet streak. Within the time frame of the COLs considered in this study,

the pattern of  $P_m$ ,  $P_e$  conversion terms (not shown), represented by the third term on the right hand side of Eq. 5, coincides with those of  $\partial_t P_e$  but with opposite sign. However the combined effect of the two terms (not shown)

$$-\nabla \cdot (\mathbf{V}P_e) + \frac{\alpha_m}{2\Theta_m} \frac{1}{d\hat{\Theta}/dp} (\mathbf{v}\theta \cdot \nabla_p \Theta_m) \quad (10)$$

209 produces composite patterns that are similar to those of  $-\nabla \cdot (\mathbf{V}P_e)$ , that are shown  
 210 in Fig. 2 (a) to (e). This means that the  $P_e$  flux divergence is the more dominant  
 211 forcing between the two, and effects the eastward propagation of the energy centre in  
 212 Fig. 1. The fluxes (represented by the arrows in the left panels of Fig. 2) show that  
 213 the energy is transported from the rear end of the potential energy centre (where there  
 214 is flux divergence) to the front end of the structure (where there is flux convergence).

215 The strength and direction of the fluxes is influenced by two factors. The first is  
 216 the amount of  $P_e$  in the jet streak, and the strength of zonal flow of the jet streak  
 217 itself. The second factor is the direction of the meridional flow, which is poleward  
 218 (equatorward) at the jet entrance (exit) region, but also weaker than the zonal jet  
 219 streak flow (Ndarana et al., 2020). This causes the fluxes to be strong in the  $P_e$   
 220 and weaker everywhere else, and further highlights the reason why the  $P_e$  structure  
 221 follows the jet streak.

222 The process that links  $P_e$  to  $K_e$  is baroclinic conversion (Orlanski and Katzfey,  
 223 1992; Orlanski and Sheldon, 1993; McLay and Martin, 2002; Decker and Martin,  
 224 2005; Harr and Dea, 2009). In Eqs 5 and 7, this process is represented by the term  
 225  $\omega\alpha$ . When  $\omega\alpha < 0$ , then a conversion from  $P_e$  to  $K_e$  takes place. Composites of  $\omega\alpha$   
 226 shown in Fig. 2 (f) to (j) demonstrate that baroclinic conversion during the evolution  
 227 of South African COLs dominates during the six hourly time steps leading up to the  
 228 time step at which the systems form at  $t = 0$  hours. These composites also show that  
 229 it occurs in the rear end of the  $P_e$  centre, represented in Fig. 2 by the thick solid  
 230 contour and is associated with increasing midlatitude baroclinicity as demonstrated  
 231 by the increasing strength of the jet streak (Ndarana et al., 2020). Note that the  
 232 relative location of  $\omega\alpha < 0$  and  $-\nabla \cdot (\mathbf{V}P_e) < 0$  means that  $\partial_t P_e < 0$  found in

233 the rear end of the  $P_e$  centre is caused by both baroclinic conversion and  $P_e$  flux  
234 divergence, with the former ceasing earlier and the latter continuing beyond the day  
235 of COL formation.

### 236 *3.2. Downstream development during COL evolution*

237 We now discuss the evolution of  $K_e$  (Fig. 3 (a) to (e)) and contrast it with that  
238 of  $P_e$  (Fig. 1 (a) to (e)) - thus making the case for downstream development that is  
239 associated with COL pressure systems. The main difference between the structures  
240 of the two energy forms is that the former has two centres; one (the midlatitude  $K_e$   
241 centre) located in the confluent region of the midlatitude jet streak and another (re-  
242 ferred to as the subtropical  $K_e$  centre) develops north east of jet streak, where we find  
243 anticyclonic barotropic shear and PV overturning. This places the former upstream  
244 and the latter downstream prior to formation of the closed circulation in COLs. We  
245 will therefore use the terms "midlatitude  $K_e$  centre" and "upstream  $K_e$  centre" as  
246 well as "subtropical  $K_e$  centre" and "downstream  $K_e$  centre" interchangeably. In  
247 addition, at each time step,  $P_e$  has only one centre, always located east (upstream)  
248 of the midlatitude  $K_e$  centre. After developing, the  $K_e$  centres maximize at different  
249 times during the evolution of COLs, with the midlatitude  $K_e$  doing so first.

250 When the  $K_e$  structures are viewed relative to the jet streak and the closed COL  
251 circulation north east of it, they may be seen as two separate entities, because whilst  
252 the jet streak influences the COLs, the two are different processes. However, when  
253 viewed relative to the composite RWB, which highlights the ridge/trough/ridge sys-  
254 tem at play, they may be seen as a clear case of downstream development, as described  
255 in Orlanski and Sheldon (1995), (cf. their Fig 3). From this point of view, Stage 1  
256 occurs at around  $t = -36$  hours in Fig. 3 (a), when the midlatitude  $K_e$  (which corre-  
257 sponds to the western centre in Orlanski and Sheldon (1995)) propagates east, whilst  
258 increasing in strength and the upstream centre develops. The saturation of this up-  
259 stream midlatitude energy centre, whilst the one downstream continues to intensify  
260 and approaching its maximum, is a clear signature of stage 2 in a developing baroclinic  
261 wave that occurs in Fig. 3(c). Stage 3 occurs at  $t = 0$  hours (Fig. 3(d)), because the  
262 western centre has begun to dissipate, whilst the downstream reaches its maximum.

263 Beyond stage 3, both energy centres dissipate. Composite evolution of COLs of all  
264 durations were examined and they exhibit exactly the same behaviour. It follows  
265 then that the COLs in the South African domain are preceded by the downstream  
266 development of a baroclinic wave.

267 Similarly to Orlanski and Sheldon (1995), downstream  $K_e$  saturates at higher  
268 values than the one up stream. There are, however, important differences between  
269 the observations described above and the Orlanski and Sheldon model of downstream  
270 development. First the upstream centre is much larger than the one located down-  
271 stream, which is consistent with the Gina and Piva (2013) COL case. Secondly, the  
272 downstream centre appears to be quasi-stationary, relative to the upstream structure  
273 - and the former eventually actually moves past as dissipation occurs. As a result  
274 of the relative speed of the two centres, the baroclinic wave has a southwest/northeast  
275 orientation so that the trough axis has a northwest/southeast slant, as opposed to  
276 that of the Orlanski and Sheldon idealised model, in which the trough axis is parallel  
277 to the latitude axis.

278 To enable a direct comparison with the diagnostics employed by Tennant and  
279 Reason (2005), the baroclinic kinetic energy is shown in Fig. 3 (f) to (j). In the  
280 midlatitudes, both forms of kinetic energy are placed at more or less the same position  
281 relative to the jet core (i.e. in the confluent region of the jet). They also propagate  
282 eastward in unison with the streak and saturate at the same time (before the formation  
283 of the COL). The difference between the two, though, is that the subtropical energy  
284 centre is placed north east of the closed circulation and develops as the small scale jet  
285 streak does the same, during the formation of the split jet (Ndarana and Waugh, 2010;  
286 Reyers and Shao, 2019; Ndarana et al., 2020). Both centres in the case of baroclinic  
287 kinetic energy can thus be considered an artefact of jet streaks, which stands to reason  
288 since these are regions of strong low level meridional temperature gradients.

### 289 *3.3. The generation and movement of the midlatitude eddy kinetic energy*

290 The use of Eq. 7, processes that inform the evolution of the two  $K_e$  centres will  
291 now be explained, by first considering the  $\partial_t K_e$  fields. These are the shaded areas in  
292 Fig. 4 (a) to (e). As was the case with  $\partial_t Pe$ , the distribution of  $\partial_t K_e$  relative to  $K_e$

293 centre is such that  $\partial_t K_e < 0$  and  $\partial_t K_e > 0$  are found at the rear and front ends of the  
 294 approximately oval  $K_e$  shape, respectively. This ensures that the maximum values  
 295 of  $K_e$ , with respect to time, occur where  $\partial_t K_e = 0$ , which divides the  $K_e$  diagonally  
 296 across, from the north-west to the south-east. These source and sink regions of  $K_e$   
 297 are consistent with the idealised model of Orlanski and Sheldon (1995).

298 The source region ahead of the midlatitude  $K_e$  centre is, first and foremost, as-  
 299 sociated with the baroclinic conversion of  $P_e$  to  $K_e$ . Comparing the corresponding  
 300 panels in Figs 2 (f) to (j) and 4 (a) to (e) shows that regions of  $-\omega\alpha > 0$  cover regions  
 301 of  $\partial_t K_e > 0$  that are slightly upstream. It is important to note at this point that  
 302  $K_e(P_e)$  is located in the confluent (diffluent) region of the jet streak. This relative  
 303 position of the energy forms makes sense because energy conversion occurs from  $P_e$   
 304 to  $K_e$ . The baroclinic conversion in the PV overturning region is weak. It is thus of  
 305 no significance and will not be discussed further.

306 The source region of the midlatitude  $K_e$  centre is secondly impacted or informed  
 307 by the  $K_e$  flux convergence (i.e.  $-\nabla \cdot \mathbf{V}K_e > 0$ ). These are shown in Fig. 4 (f) to (j).  
 308 This field affects  $\partial_t K_e > 0$  slightly downstream of the region where  $P_e$  is converted  
 309 to  $K_e$ . The rear end of the  $K_e$  centre is characterised by divergence and the flux  
 310 vectors,  $\mathbf{V}K_e$ , are oriented consistently to this. The fluxes that move energy from  
 311 the back to the front of the energy centre are of significant size in the middle of the  
 312 centre, for similar reasons to those that were discussed in Subsection 3.1. Arguments  
 313 as to how the  $K_e$  translates eastward are, thus, similar to those of  $P_e$ . Outside the jet  
 314 streak, in the PV overturning region (close to the closed COL circulation), the eddy  
 315 kinetic fluxes are oriented in a northward direction and their convergence is mostly  
 316 concentrated in the upstream end of the subtropical energy centre (eastern energy  
 317 centre). This phenomenon is most clearly seen from about  $t = -24$  to 0 hours in Fig.  
 318 4 (g) to (i). The fact that the energy flux is stronger in the western half of this  $K_e$   
 319 centre explains why it is oriented as it is. Its rear (or upstream) end moves further  
 320 northward as compared to the eastern part. The orientation of the fluxes in the PV  
 321 overturning region is caused by the fact that the zonal flow is significantly decelerated  
 322 there to values close to zero and is, in some cases, slightly negative (Ndarana et al.,

323 2020). The direction of the fluxes is consistent with the flux divergence (convergence)  
324 that is found on the south (north) end of the subtropical  $K_e$  centre.

### 325 *3.4. The evolution of the subtropical eddy kinetic energy centre*

326 As noted above, baroclinic conversion in the PV overturning region (where the  
327 subtropical  $K_e$  centre is located, highlighted by the green box in Fig. 5) is small (see  
328 Fig. 2 (f) to (j)) and secondly, again as noted above, the  $K_e$  fluxes move the centres.  
329 It follows then that these two processes cannot be responsible for the generation and  
330 growth of the subtropical  $K_e$  centre. Instead, its growth comes from receiving energy  
331 by means of the ageostrophic geopotential fluxes (left panels Eq. 5) which "radiate"  
332 energy from the front end of the midlatitude  $K_e$  into its rear end. The exact area in  
333 the midlatitude centre where the energy originates is indicated by  $\nabla_p \cdot (\mathbf{v}\phi)_a > 0$ . The  
334 notion of radiative energy transfer is used here to distinguish this process from the  
335 fluxes (Orlanski and Sheldon, 1995) that were discussed in the previous section, as the  
336 two types of fluxes play different roles. These radiative energy transfer processes start  
337 developing from  $t = -48$  hours (not shown), progressively evolve and mature through  $t$   
338  $= -12$  hours as they curve more sharply (as the red contour shows the deepening ridge).  
339 They appear to reach their maximum strength when the subtropical  $K_e$  reaches its  
340 maximum on  $t = +0$  hours (Fig. 5 (d)), after which it wanes. The downstream energy  
341 centre thus develops and grows by receiving energy from the upstream midlatitude  
342 centre via  $(\mathbf{v}\phi)_a$ .

343 Simple physical arguments can be used to explain the onset of ageostrophic fluxes,  
344 their increase in strength and orientation. The area between the two energy centres  
345 (green box) in Fig. 5, characterised also by positive PV anomalies (i.e. the western  
346 lobe of the breaking wave), is a ridge and therefore the flow in the region is anticyclonic  
347 and supergeostrophic (Lim and Wallace 1991). Therefore, the ageostrophic circulation  
348 would also exhibit anticyclonic behaviour (see orientation of the arrows in the left  
349 panels of Fig. 5 (f) to (j)). It is important to note that the subgeostrophic nature of  
350 the flow at the top of the trough, above the area of the closed circulation. This is  
351 entirely consistent with our understanding of the behaviour of flow, thus confirming  
352 our findings.



353 Since the highlighted area is a ridge, it follows that  $\phi > 0$  (brown shading in the  
354 left panels of Fig. 5 (f) to (j)). The flux vectors will have the same direction as the  
355 ageostrophic flow and will therefore be directed as shown in the area highlighted by  
356 the green box, particularly from  $t = -24$  hours (Fig. 5 (g)) to  $t = +12$  hours (Fig.  
357 5 (j)). As the waves break, the values of  $\phi$  increase in magnitude. This is clearly  
358 an effect of the deepening ridge, as tropospheric air is advected anticyclonically and  
359 poleward by the wave breaking processes. Therefore the combined effect of increase  
360 in the strength of  $\mathbf{v}_a$ , its direction and the increasing magnitude of  $\phi$  explains the  
361 development and increasing strength of  $(\mathbf{v}\phi)_a$ .

### 362 3.5. Linkages of COLs to midlatitude processes

363 Downstream development provides a framework of linking COLs that impact  
364 South Africa to midlatitude processes. Given that, we now consider the relationship  
365 between the diagnostics presented above in a geographical context of South Africa  
366 and attempt to link the results obtained here to those of Tennant and Reason (2005).

367 Following Singleton and Reason (2007a), we divide the South African domain (see  
368 Subsection 2.2) into four subdomains (see Fig 6 (a)), A (red), B (blue), C (green), and  
369 D (magenta), that are bounded by  $(10 - 27^\circ\text{E}, 30 - 40^\circ\text{S})$ ,  $(10 - 27^\circ\text{E}, 20 - 30^\circ\text{S})$ ,  $(27 -$   
370  $40^\circ\text{E}, 20 - 30^\circ\text{S})$  and  $(27 - 40^\circ\text{E}, 30 - 40^\circ\text{S})$ , respectively. Unlike Singleton and Reason  
371 (2007), though, we say that a COL event belongs to the region in which it was first  
372 identified, even if it evolves into the downstream (or upstream for that matter). Of  
373 the 476 COL cases that were identified in this study, 232 (45%), 107 (21%), 52 (10%)  
374 85 (17%) are found in regions A, B, C and D, respectively. All four regions exhibit a  
375 minimum number of COLs during the summer and maximum occurs in October for  
376 region A and B, in April for region D and in May for region C.

377 We now make a case for the notion that the  $K_e$  associated with COLs that affect  
378 South Africa originates in the midlatitudes. As discussed in Subsection 3.4, the  
379 midlatitude  $K_e$  propagates eastward with the jet, and as it does, energy is transferred  
380 from it by means of ageostrophic geopotential fluxes so that the downstream  $K_e$   
381 grows, in strength. In this section we present this result in a geographical setting  
382 to show that depending on where the COLs form in the South African domain, the

383 downstream  $K_e$  centre will behave differently. For Region A COLs (left panels of  
384 Fig. 7) and Region B COLs (right panels of Fig. 7) the downstream  $K_e$  centre  
385 develops west of 20°E in the South Atlantic Ocean. For both categories of COLs, it  
386 remains west of this latitude line and then propagates towards the southwestern tip of  
387 Africa. At  $t = 0$  hours, the  $K_e$  associated with Region A COLs is eventually located  
388 south of the one for Region B and it is stronger. The energy transfer occurs as the  
389 midlatitude jet streak propagates eastward and increases in strength, thus bringing  
390 with it increasing anticyclonic barotropic shear, and increasing strain rate (Nakamura  
391 and Plumb 1994) that leads wave breaking on the 330 K dynamical tropopause in the  
392 case of Region A COLs. The wave breaking induces the ageostrophic geopotential  
393 fluxes, as shown in Subsection 3.4. Simple experimentation shows that for Region B  
394 COLs, the wave breaking that influences the fluxes is most clearly seen on the 340 K  
395 isentropic surface. The associated of Region A and B COLs with RWB on the 330  
396 K and 350 K, respectively, explains the  $t = 0$  hours position (Figs 7 (d) and (i)) of  
397 the respective  $K_e$  centres, relative to one another. The difference in the intensity of  
398 the eddy kinetic energy density of the two categories of COLs can also be explained  
399 by the RWB. Observing that the  $PV = -1.5$  PVU contour turns back on itself in  
400 the case of Region A and not in the case of Region B COLs shows that the fluxes  
401 associated with the latter will be weaker and hence the eddy kinetic energy associated  
402 with them.

403 As noted in the Introduction, Tennant and Reason (2005) found that wet South  
404 African seasons are associated with two branches of the baroclinic kinetic energy.  
405 Fig. 7 shows this split (thin black contours). This figure also shows that COLs that  
406 develop in the western half of South Africa are associated with a large scale baroclinic  
407 kinetic energy structure that is located over subtropical South Atlantic, South African  
408 mainland and South West Indian Ocean, and oriented in a northwest/southeast slant.  
409 It is quasi-stationary relative to the one observed to appear to be moving with the  
410 midlatitude jet streak.

411 Fig. 7 suggests that the subtropical baroclinic kinetic energy centre is influenced  
412 by midlatitude processes. Even though, to the best of our knowledge, there is no

413 known theoretical framework that unifies the energy equation systems of Wiin-Nielson  
414 (1962) and Orlanski and Katzfey (1991), Fig. 7 suggests that the ageostrophic geopo-  
415 tential fluxes might be responsible for the variations of the subtropical baroclinic ki-  
416 netic energy. The magnitude of the subtropical kinetic increases (decreases) as the  
417 fluxes strengthen (weaken) during wave breaking. It appears as though it is influenced  
418 by the fluxes that are influenced by the subgeostrophic flow from the west over the  
419 South Atlantic Ocean, which are associated with the presence of COLs and RWB in  
420 that region. Also the supergeostrophic flow from the south-west (and the increasing  
421  $\phi > 0$  discussed in Subsection 3.4) causes the fluxes to be orientated northeastward,  
422 into the eastern half of the subtropical baroclinic kinetic energy branch. In this study,  
423 we thus propose that the branches of baroclinic kinetic energy of Tennant and Reason  
424 (2005) are connected.

425 The transfer of eddy kinetic energy from the midlatitudes into the eastern parts  
426 of the South African domain is facilitated by wave breaking on the 340 K and 330 K  
427 dynamical tropopause for Region C and D COLs, respectively. Note that the RWB  
428 processes associated with these COLs occur downstream from those shown in Fig.  
429 8 and discussed above. The  $K_e$  density associated with Region C COLs is much  
430 weaker than its Region D COL counterpart. This is caused by the much weaker  
431 ageostrophic geopotential fluxes out of the midlatitudes, as informed by the depth of  
432 the tropopause fold associated with them. Observe the behaviour of the  $PV = -1.5$   
433 PVU (thick red contour), which is much more deformed in the case of Region D COLs  
434 than in that of Region C COLs. When the deformation of this contour is compared  
435 across all categories of COLs (Figs 7(d), (i), 8(d) and (i)), it becomes apparent that  
436 the depth of the PV anomaly might be playing a role in the strength of the  $K_e$  and  
437 might related to the extension of the COLs to the surface (see Barnes et al., 2020).  
438 This will be a subject of further analysis because it is beyond the scope of the current  
439 study. The weak nature of the Region C COL  $K_e$  density and orientation of the  
440 associated fluxes means that the COLs that occur in this area have little connections  
441 to the midlatitudes, except the fact that the PV anomalies that induce them are a  
442 results of wave breaking that is caused by the midlatitude jet.

443 In stark contrast to Fig. 7, inspection of Fig. 8 indicates that the subtropical  
444 baroclinic kinetic energy is located in the Indian Ocean for COLs that develop over  
445 the eastern half of the country. For Region D COLs in is more further east than  
446 for Region C COLs. Therefore, if there is any connection between the subtropical  
447 branch of baroclinic kinetic energy to its midlatitude counterpart, it happens much  
448 further downstream and not over South Africa as we have found that to be the case  
449 for western COLs. The direction of the fluxes that might influence these subtropical  
450 baroclinic energy structures are direction consistently to where they are located. The  
451 same applies to ageostrophic geopotential fluxes associated with the subgeostrophic  
452 flow leaving subcontinent from southern Mozambique.

#### 453 4. Concluding remarks

454 Using 39 years (1979 - 2018) of ECMWF reanalysis data, an established local en-  
455 ergetics framework (Orlanski and Katzfey, 1991) and simple composite analysis, this  
456 study has shown that South African COL pressure systems are preceded by down-  
457 stream development of a baroclinic wave (Orlanski and Sheldon, 1995). This process  
458 is most clearly seen by examining eddy kinetic energy ( $K_e$ ), which is converted from  
459 eddy available potential energy ( $P_e$ ) as found in classical energetics theory (Lorenz,  
460 1955). However additional processes are required to be considered to complete the  
461 picture that emerges in a spatially varying setting. To summarise the processes in-  
462 volved, we proceed as follows:

- 463 1. A few days before the formation of the closed COL circulation, the midlatitude  
464 jet streak first propagates in the south eastward direction (by means of momen-  
465 tum advection processes - Ndarana et al., 2020) and then more zonally, whilst  
466 gaining in strength. The jet streak propagates together with  $P_e$  in its diffluence  
467 regions and  $K_e$  further upstream in the confluence of the streak.
- 468 2. This midlatutide  $K_e$  centre grows by gaining energy from the  $P_e$  ahead of it  
469 by means of baroclinic conversion, which continues up to the point when the  
470 closed COL circulation forms and appears to cease thereafter as the jet streaks  
471 passes south of the COLs. The movement of the  $K_e$  is caused by energy fluxes

472 by the total flow within the energy centre and they distribute the energy from  
473 the rear to front end of the centre. The strength and direction of the fluxes  
474 are influenced by the flow of the jet streak in the case of the midlatitude  $K_e$   
475 centre.

- 476 3. The propagation of the jet streak and its increasing zonal flow, coupled with the  
477 smaller scale jet streak north of the COL region, constitute a split jet found in  
478 previous studies (Ndarana and Waugh, 2010; Reyers and Shao, 2019), which in  
479 turn, increases anticyclonic barotropic shear and shearing strain (Nakamura and  
480 Plumb 1994) leading to anticyclonic RWB (Peters and Waugh, 2003), signalled  
481 by PV overturning. The wave breaking processes create a ridge southwest of  
482 the COL circulation but on the equatorward side the jet and this ridge deepens  
483 as wave breaking evolves. As a result the flow becomes increasingly super-  
484 geostrophic and the geopotential anomalies deepen, thus inducing ageostrophic  
485 geopotential fluxes, and with time, increasing their magnitude. The super-  
486 geostrophic flow that is associated with the wave breaking is directed anticy-  
487 clonically, which in turn informs the direction of the fluxes towards the COL  
488 regions because the geopotential perturbations are positive.
- 489 4. These ageostrophic geopotential fluxes are responsible for transferring energy  
490 from the upstream  $K_e$  to the one downstream. Thus the latter grows, not from  
491 baroclinic conversion, but from ageostrophic geopotential flux convergence. It  
492 then reaches a maximum at the point when the closed COL circulation forms.

493 Overall, the upstream  $K_e$  centre increases, whilst another  $K_e$  centre develops  
494 downstream. The former reaches its maximum before the formation of the COLs,  
495 and the latter continues to grow. By the time the downstream structure reaches a  
496 maximum at the point in time that the COLs form, the centre upstream has begun  
497 to dissipate. This is a sequence of events that characterises downstream development  
498 and the growth and decay of these energy centres are informed by the processes listed  
499 above. Therefore COLs do indeed develop from unstable synoptic scale Rossby waves  
500 (Favre et al, 2012; Omar and Obiodun, 2020). This is also demonstrated by PV,  
501 which evolves to the point of over turning of the PV contour.

502 This study has shown that COLs that occur over the northwestern, southwestern  
503 and southeastern parts of the South African domain, as defined, here have strong  
504 links to the midlatitudes. Using the downstream development framework, it was  
505 shown that the  $K_e$  that is associated with COLs is transferred from the midlatitude  
506 by energy radiative processes that are induced and enhanced by wave breaking on the  
507 midlatitude dynamical tropopause on different isentropic surfaces. These midlatitude  
508 connections during the evolution of dynamical processes leading up to COL formation  
509 also appear to connect the subtropical branch of the baroclinic kinetic energy to its  
510 midlatitude counter part that were identified by Tennant and Reason (2005). It was  
511 shown in that study that when configured in this way, the two centres of baroclinic  
512 eddy kinetic energy characterise wet South African seasons. It should be mentioned,  
513 however, while the connection between the two baroclinic kinetic energy is clear, it is  
514 necessary to quantify it in a future investigation.

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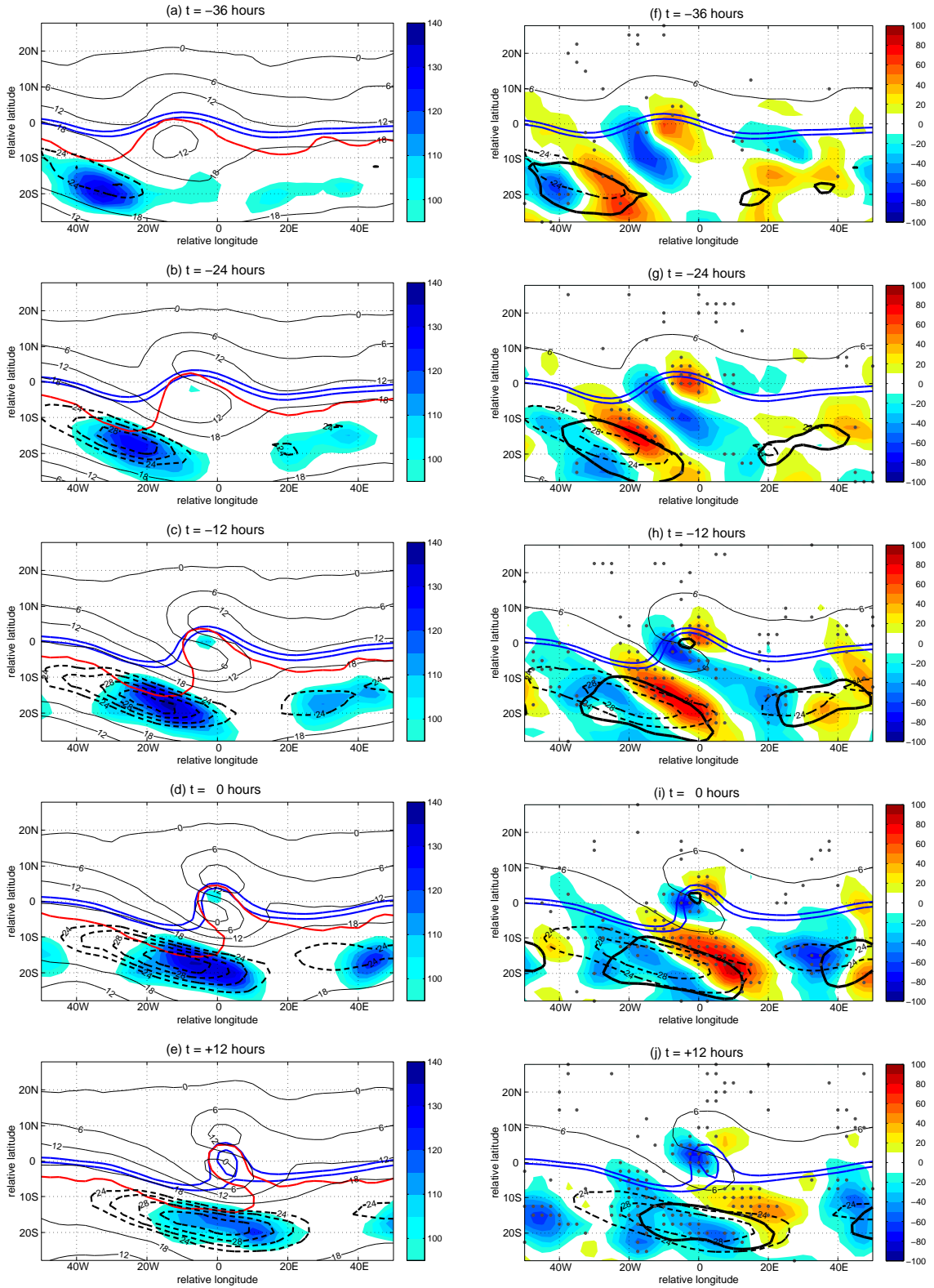


Figure 1: Time-lagged composites of vertically integrated eddy available potential energy (shaded in blue) shown on the left panels. The eddy available potential energy is plotted from 95 - 140  $\text{m}^2 \text{s}^{-2}$ . The thick red contour is the -2 PVU contour on the 330 K isentropic surface. The right panels show time-lagged composites of the tendency of vertically averaged eddy available potential energy (shaded) plotted in  $\text{m}^2 \text{s}^{-2} \text{day}^{-1}$ . In all panels, the thick blue contours are the 5614 and 5635 gpm geopotential height. The thin black contour is the  $6 \text{ m s}^{-1}$  zonal isotach and the thick dashed black contour is the 24 to  $28 \text{ m s}^{-1}$  zonal isotach. The thick dashed black solid contour in the right panels is the  $100 \text{ m}^2 \text{s}^{-2}$  contour of eddy available potential energy and the grey dots represent areas where the tendency of the eddy available potential energy is significant at the 90% level. The composites are plotted in 12 hour intervals from (a,f)  $t = -36$  hours to (e,j)  $t = +12$  hours.

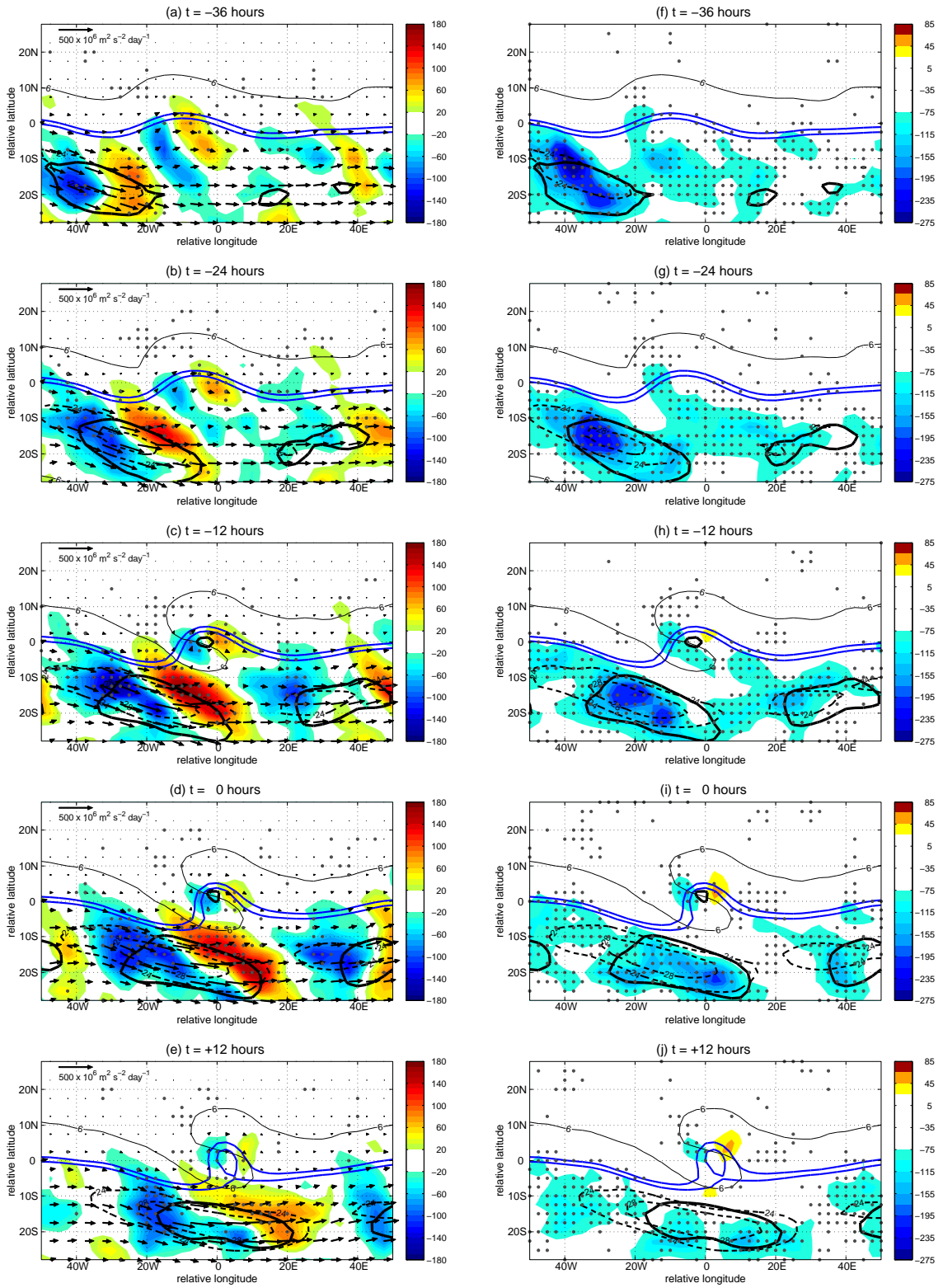


Figure 2: Time-lagged composites of eddy available potential energy fluxes  $\mathbf{VP}_e$  (arrows) and their divergence  $-\nabla \cdot (\mathbf{VP}_e)$ , shaded) in left panels and baroclinic conversion  $(\omega\alpha)$  shaded in the right panels, both plotted in  $\text{m}^2 \text{s}^{-2} \text{day}^{-1}$ . The blue contours, thin and thick dashed black contours and thick solid black contour and evolution time steps as in Fig. 1.

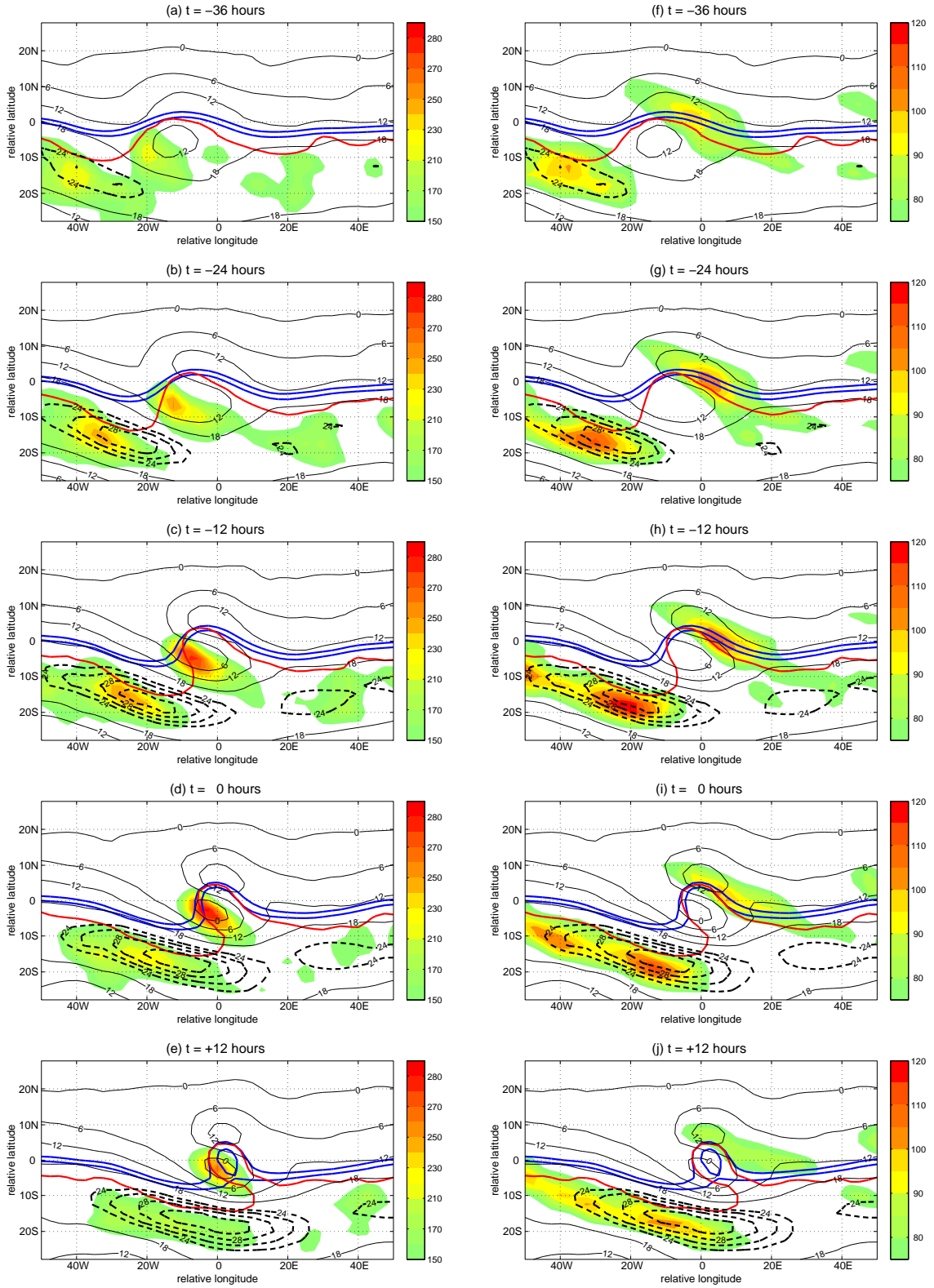


Figure 3: Time-lagged composites of  $K_e$  (shading on the left panels) and  $K_{BC}$  (shading on the right panels). The energy fields are plotted in  $\text{m}^2 \text{s}^{-2}$ . The blue contours, thin and thick dashed black contours, thick solid black contour and thick red contour and evolution time steps as in Fig. 1.

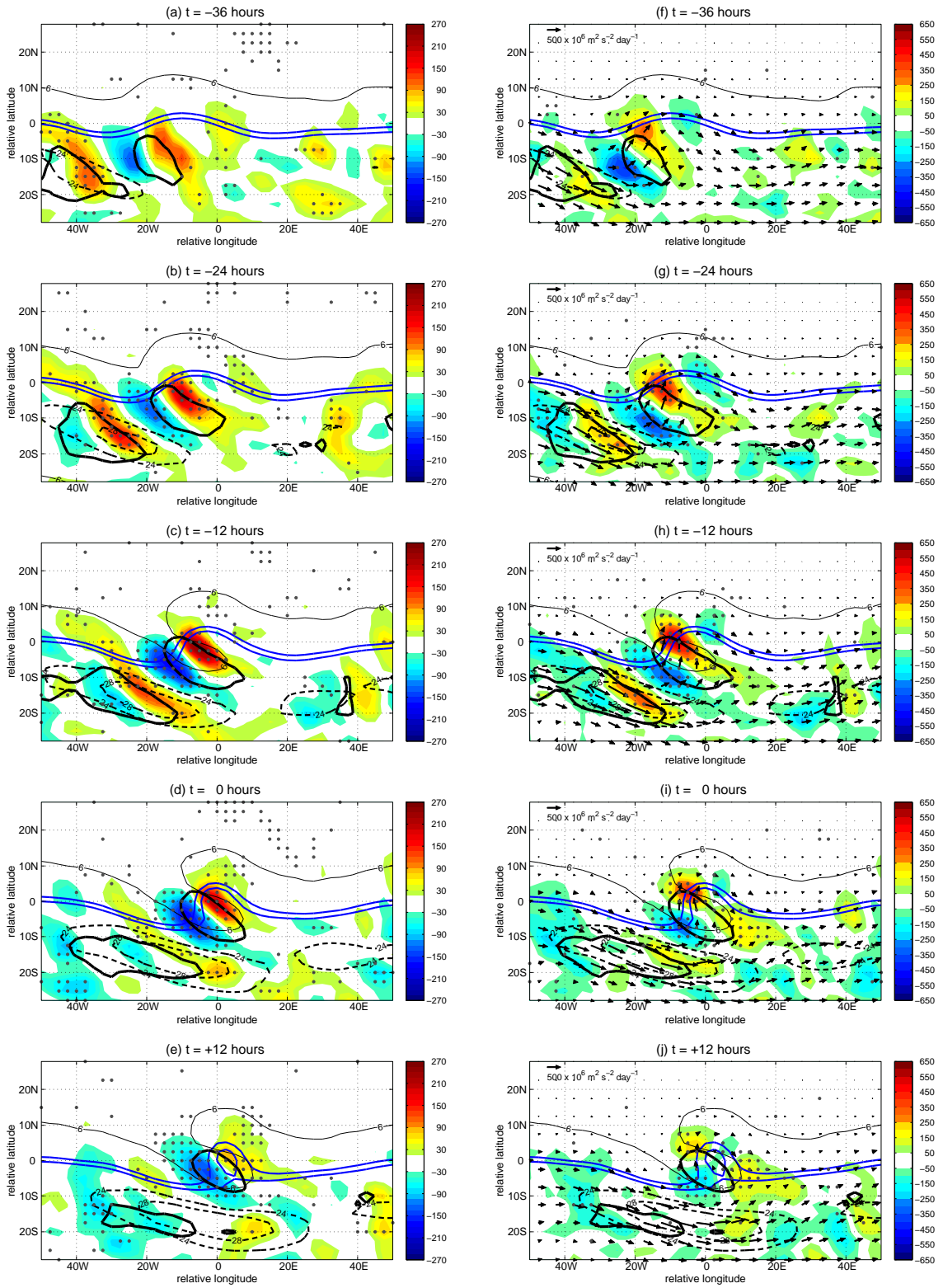


Figure 4: Time-lagged composites of  $\partial_t K_e$  (shading on the left panels) and  $\nabla \cdot (\mathbf{V}K_e)$  (shading on the right panels) and flux vectors  $\mathbf{V}K_e$ . The  $K_e$  tendency and the flux divergence are plotted in  $\text{m}^2 \text{ s}^{-2} \text{ day}^{-1}$ . The thick solid black contour is the  $170 \text{ m}^2 \text{ s}^{-2}$ . The blue contours, thin and thick dashed black contours and evolution time steps as in Fig. 1.

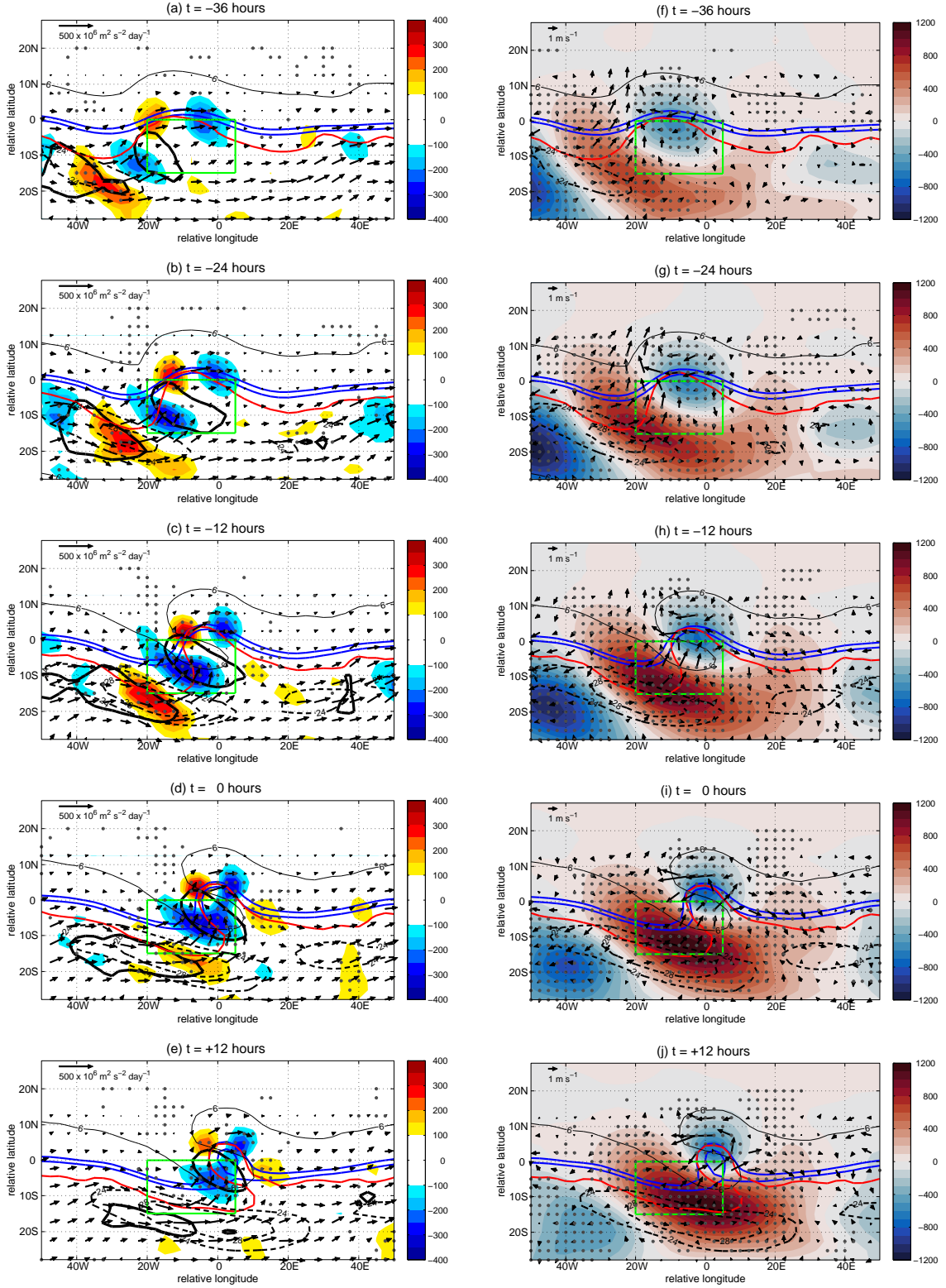


Figure 5: Time-lagged composites of  $-\nabla \cdot \mathbf{v}_a \phi$  (shading on the left panels) and the arrows represent the ageostrophic geopotential flux  $\mathbf{v}_a \phi = (\mathbf{v} - f^{-1} \mathbf{k} \times \nabla \phi) \phi$ . The shading in the right panels represent  $\phi$  and the arrows represent the ageostrophic flow  $\mathbf{v}_a = u_a \mathbf{i} + v_a \mathbf{j}$ . The thick solid black contour is the  $170 \text{ m}^2 \text{ s}^{-2}$ . The blue contours, thin and thick dashed black contours, and thick red contour and evolution time steps as in Fig. 1.



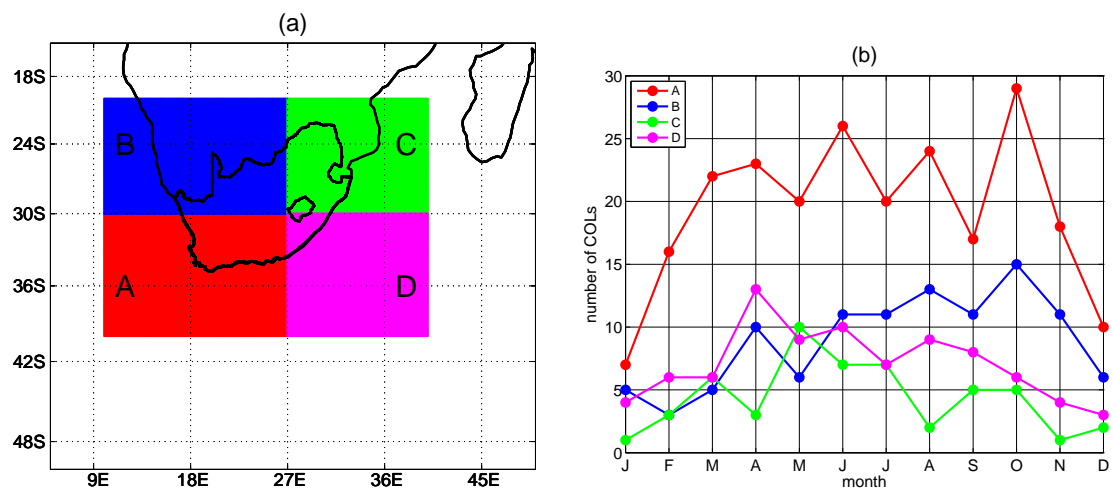


Figure 6: (a) South Africa divided into climatological regions A, B, C and D for COLs (Adapted from Singleton and Reason (2007)). (b) Monthly variations of COLs for the regions in (a).

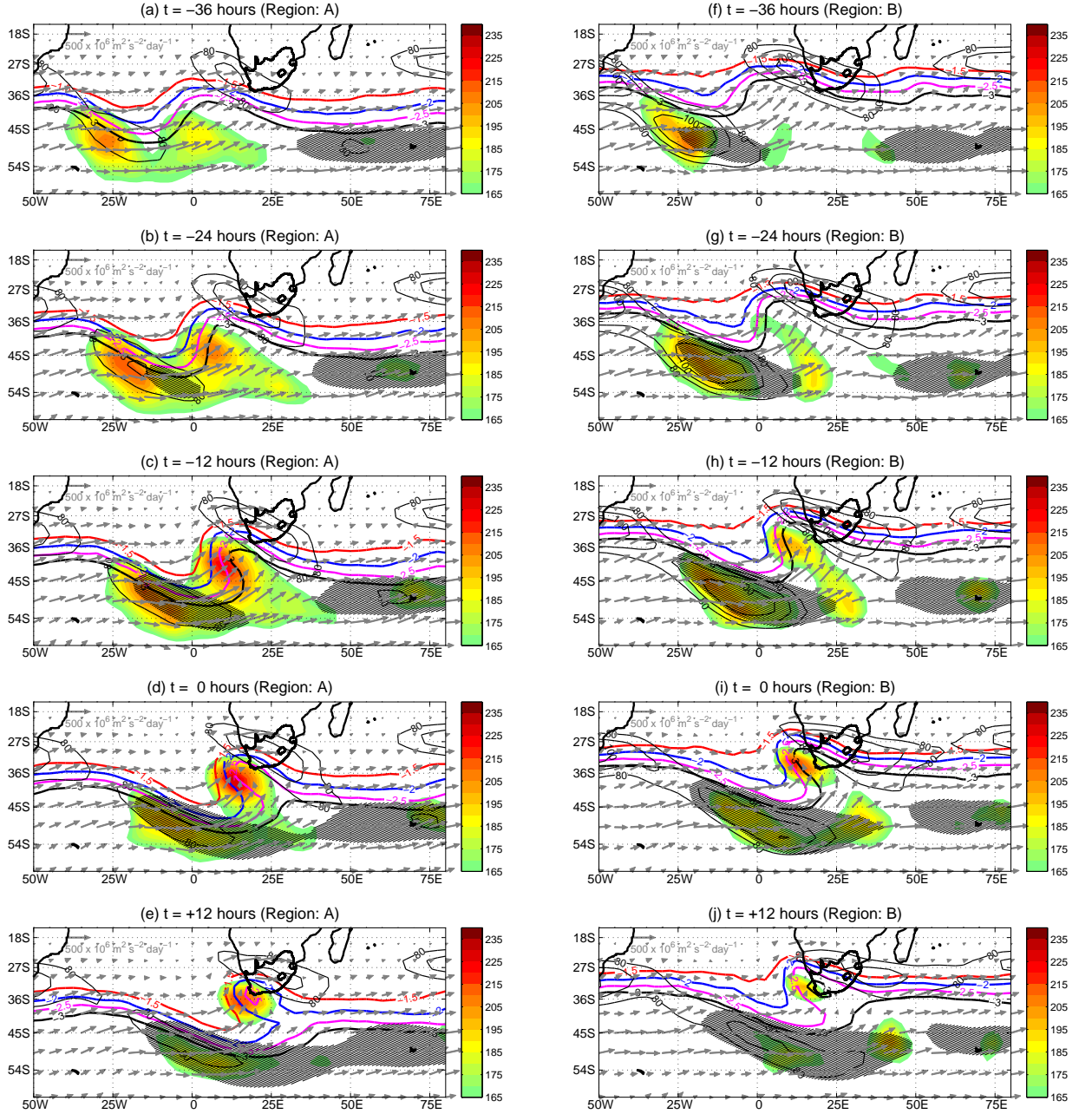


Figure 7: Time-lagged composites of eddy kinetic energy (dashed), zonal wind ( $25 \text{ m s}^{-1}$  or greater hatched) and baroclinic kinetic energy (thin black contours). The kinetic energy is plotted in  $\text{m}^2 \text{ s}^{-2}$ . The thick red, blue, magenta and black contours represent the -1.5, -2, -2.5 and -3 PVU contours, respectively, on the 330 K (left panels) and 340 K (right panels) isentropic surfaces. The arrows represent the ageostrophic geopotential fluxes as in Fig. 5. The left (right) panels are composites created using COLs whose initial points are in region A (B). Composites are plotted from (a, f)  $t = -36$  hours to (e, j)  $t = +12$  hours.

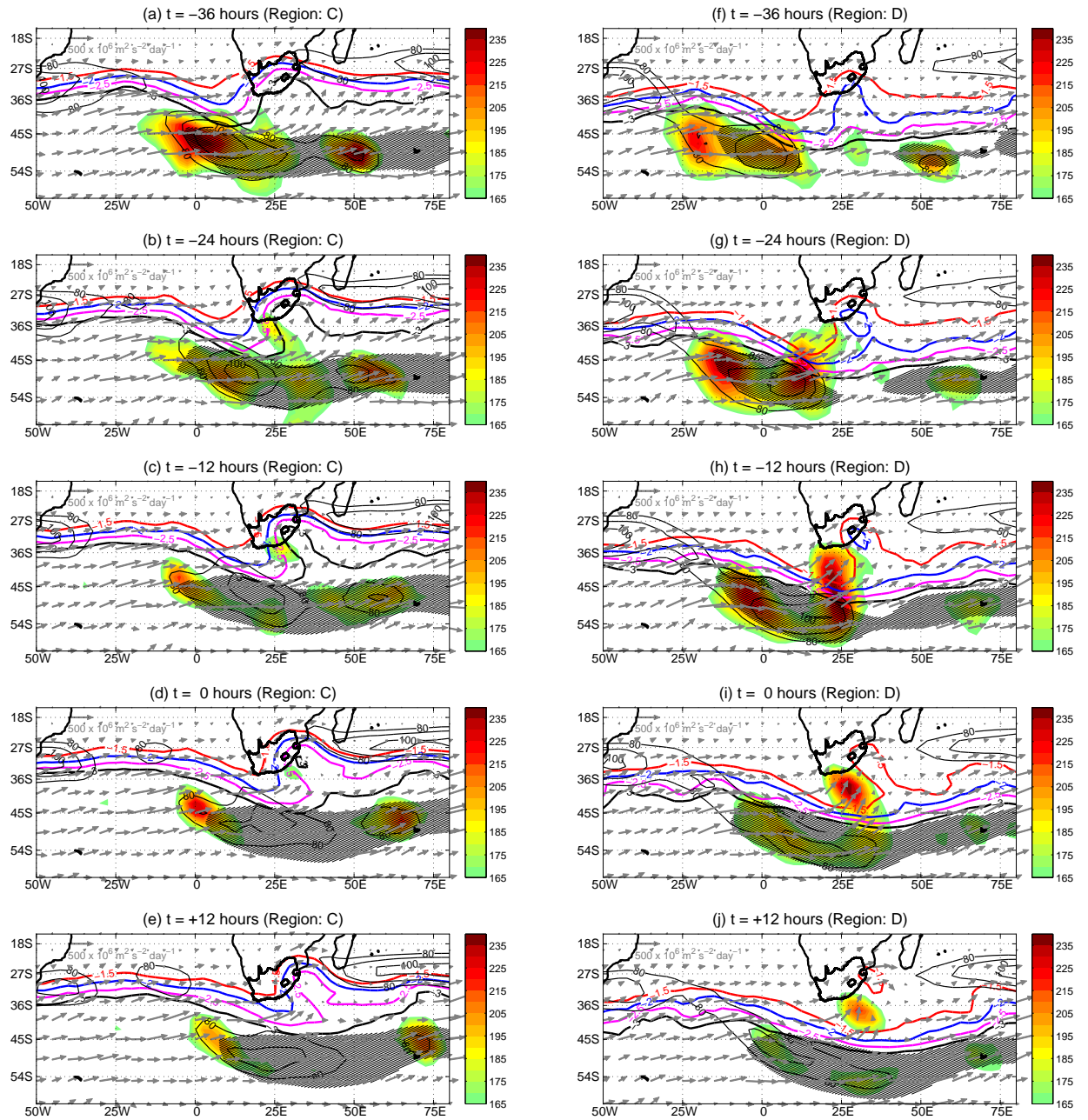


Figure 8: Same as Fig. 7 but for region C (D) on the left (right) panels. The PV on the left (right) panels is on the 340 (330) K isentropic surface.